

KODAIKANAL OBSERVATORY

BULLETIN Number 181

Photoelectric light curve of YY Eridani

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Abstract

Photoelectric light curves in B and V of the W Ursa Majoris star YY Eridani have been determined. Gradual lengthening of the period has been observed.

The light curve of the star YY Eridani, HD 26609, an eclipsing binary was obtained by a photoelectric photometer employing an RCA 1P21 photomultiplier tube, on twelve nights during December 1965—March, 1966 at Kodaikanal. The telescope used was a 20cm refractor. The signal from the photomultiplier was amplified by a linear D.C. amplifier and recorded on a potentiometric type stripchart recorder. The observations were taken in two colours, blue and yellow using standard glass filters of the UBV system with pass bands centred around 4500Å and 5500Å. For comparison, two other nearby stars were observed. The details of the three stars are given below:

Star	HD No.	(1950)	(1950)	Spectral type
YY Eridani	26609	1h 9m 16s.125	- 10° 35' 19" .58	G5, G7
Comparison	26650	1h 10m 9s.980	- 10° 41' 35" .20	A5
Check	26902	1h 12m 25s.410	- 10° 11' 36" .56	K0

Altogether 468 comparisons in blue and 422 in yellow light were obtained during this period. Six primary and three secondary minima were observed. It is found that to fit our observation with earlier ones by Cylie (1951) Haruhuta (1953), Kwee (1958) and Purgathofer (1961), a further change in the period must be assumed.

The new period is found to be near the value obtained by Purgathofer. The elements are well represented by the equation

$$\text{Min.} = \text{JD } 2433617^{\text{d}} \cdot 5197 \cdot 10^{\text{d}} \cdot 32149630\text{E}$$

the epoch being counted from Cillie's well determined minima of JD 2433617.5197. The observed minima timings were found to be in excellent agreement to those calculated by the above equation.

The table below gives the residuals (O-C) for the six primary minima recorded.

TABLE I
Primary Minima of YY Eridani

Heliocentric J.D.	Computed Times	Residual (O-C)
2439120.2503	2439120.2501	.0001
121.1005	124.1083	.0002
165.2596	165.2590	- .0003
106.2241	106.2213	- .0012
167.1816	167.1808	- .0002
187.1216	187.1216	.0000

The times of the three secondary minima were also determined and are tabulated in Table II. It is seen that the mean time interval between the secondary minima and the preceding primary minima is $0^{\text{d}} \cdot 1611$ which is $0^{\text{d}} \cdot 0003$ later than the midpoint of the cycle. This value is within the limits of error of the measurement and hence the secondary minima may be considered equidistant from the two nearby primary minima.

TABLE II
Secondary Minima of YY Eridani

Heliocentric J.D.	Computed time of previous primary min.	Residual (O-C)
2439162.2058	2439162.0449	.1609
165.0999	164.9301	.1615
181.1742	181.0132	.1610

The previous observers have noted a gradual lengthening of the period of this particular system. Our results also corroborate their findings. The results of the period change are represented in Fig. 1 where periods determined by several authors over the 33 year interval are plotted against the measured epochs. All the photoelectric observations can be fitted by a smooth curve. Considering the uncertainty in determining the actual times of minima by the photographic technique, the earlier photographic observations by Bodokia (1938) and Jensch (1934) can also be considered to be contained in the extended curve.

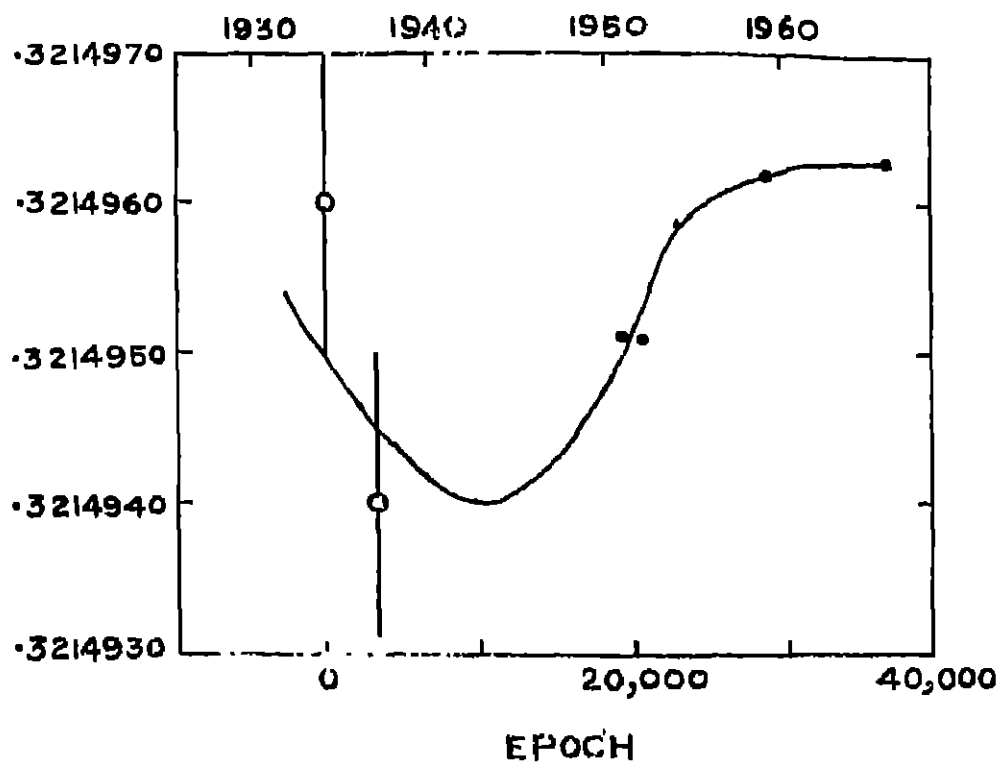


Fig. 1. Period Variation of YY Eridani

TABLE III

JD.	Δm	JD.	Δm
<i>Blue</i>			
2499120 1642	- 017*	2499120.9253	- .180
.1666	.125	.9260	.215
.1677	.125	.9300	.116
.1712	.087	.9323	.169
.1788	.169	.9397	.121
.1808	.185	.9350	.164
.1933	.126	.9370	.183
.1955	.093	.9387	.168
.2024	+ 007	.8431	.196
.2051	.015	.9161	.123
.2128	.125	.9189	.175
.2156	.031	.9515	.062
.2219	.091	.9566	.087
.2226	.058	.8586	.082
.2246	.081	.9308	.012
.2260	.105	.9649	.032
.2337	.347	.9870	.015
.2357	.558	.9607	.099
.2371	.408	.9701	.092
.2434	.117	.9746	.025
.2462	.621	.9758	.080
.2482	.600	.9771	.111
.2515	.477	.9795	- .008
.2572	.509	2499162.0925	+ .924
.2570	.151	.0933	.140
.2607	.530	.0991	.007
.2611	.517	.1008	.107
.2635	.550	.1022	.101
.2705	.386	.1092	.068
.2725	.379	.1112	.035
.2733	.309	.1126	.116
.2777	.191	.1168	.068
.2802	.157	.1182	.078
.2829	.229	.1196	.053
.2830	.169	.1230	.106
.2892	.068	.1251	.080
.2927	.033	.1265	.079
.2911	.019	.1811	.222
.2976	- 032	.1820	.270
.2990	+ 026	.1125	.120
.3010	- 007	.1152	.077
.3032	.026	.1166	.089
.3080	.097	.1180	.120
.3128	.010	.1508	.201
.3140	.021	.1612	.267
.3169	.061	.1702	.193
.3205	.105	.1730	.236
.3226	.130	.1751	.273
.3232	.211		

*Observations on JD 2499120 were through a narrow band interference filter ($\pm 1\%$) centered round 4200Å.

TABLE III (Contd.)

J.D.	$\Delta m.$	J.D.	$\Delta m.$
2199162	.1827	2199165	.1107
.1031	.368	.1193	.106
.1897	.133	.1200	.427
.1910	.333	.1263	.981
.1921	.539	.1284	.295
.1991	.712	.1298	.289
.2001	.700	.1367	.249
.2015	.663	.1381	.180
.2022	.616	.1395	.170
.2028	.677	.1485	.118
.2035	.656	.1485	.138
.2042	.655	.1499	.120
.2050	.626	.1517	.080
.2077	.632	.1531	.005
.2084	.613	.1603	.022
.2160	.729	.1652	.030
.2167	.561	.1707	.193
.2174	.461	.1728	.122
.2230	.459	.1712	.005
.2257	.331	.1811	.093
.2270	.501	.1832	.030
.2382	.281	.1867	.013
.2389	.350	.1881	.010
.2480	.007	.1920	.036
.2507	.278	.1961	.023
.2521	.152	.1999	.074
.2618	.007	.2020	.101
.2692	.050	.2017	.112
.2709	.025	.2103	.141
2439103	.0784	.2130	.138
.0798	.475	.2392	.403
.0812	.538	.2399	.360
.0819	.531	.2401	.173
.0840	.561	.2415	.194
.0853	.592	.2436	.487
.0867	.571	.2450	.354
.0881	.637	.2470	.605
.0902	.660	.2481	.617
.0916	.613	.2503	.656
.0944	.711	.2519	.707
.0951	.733	.2533	.735
.0958	.739	.2554	.720
.0971	.735	.2575	.706
.0985	.811	.2591	.617
.1006	.745	.2609	.076
.1020	.712	.2637	.691
.1041	.690	.2651	.629
.1055	.737	.2665	.610
.1060	.652	.2706	.346
.1082	.656	.2734	.302
.1098	.622	.2755	.323
.1159	.515	2199166	.1996
.1173	.169	.1961	.205
		.1998	.423

TABLE III (Contd.)

J.D	Δm	J.D	Δm
2439166.2033	.479	2439171.1013	.767
.2040	.472	.1055	.673
.2089	.506	.1381	.162
.2110	.545	.1430	.144
.2130	.610	.1457	+.122
.2151	.612	.1561	-.003
.2165	.690	.1575	+.110
.2172	.748	.1589	.103
.2183	.812	.1645	.023
.2207	.833	.1698	.024
.2248	.783	.1700	.057
.2262	.822	.1714	.050
.2276	.788	.1750	.053
.2290	.781	.1769	.062
.2318	.732	.1880	.031
.2332	.737	.1901	.126
.2378	.610	.1929	.036
2439167.1511	.174	.1961	.078
.1518	.240	.2033	.085
.1553	.242	.2051	.080
.1587	.272	.2123	.090
.1587	.274	.2141	.110
.1608	.300	.2165	.141
.1637	.458	.2227	.107
.1664	.494	.2248	.162
.1685	.523	.2304	.176
.1698	.549	.2325	.190
.1719	.602	.2345	.240
.1747	.661	.2359	.252
.1761	.709	.2380	.350
.1773	.759	.2401	.413
.1789	.785	.2443	.414
.1802	.782	2439175.0059	.418
.1816	.800	.0094	.350
.1837	.799	.0077	.163
.1851	.833	.0098	.200
.1865	.806	.1026	+.196
.1893	.843	.1421	-.008
.1920	.788	.1435	.002
.1927	.816	.1491	+.078
.1941	.769	.1539	.030
.1955	.782	.1615	.006
.1976	.738	.1629	.093
.2018	.642	.1713	-.147
.2032	.588	.1740	.104
.2053	.537	.1851	.080
.2073	.505	.1872	+.018
.2108	.409	.1900	.027
.2122	.357	.1921	.099
.2156	.839	.1942	.211
.2198	.273	.1962	.204
.2267	.149	.1983	.206
.2281	.135	.2004	.375
.2323	.088	.2032	.439

TABLE III (Contd.)

J.D	Δm	J.D	Δm
2489175.2019	.471	2489181.0816	.023
.2112	.625	.0909	.023
.2148	.716	.0930	.027
2489176.0775	.328	.0999	.015
.0837	.083	.1019	.070
.0851	.061	.1075	.056
.0879	.059	.1103	.017
.0900	.091	.1131	.093
.0962	.098	.1145	.079
.1011	.092	.1159	.058
.1030	.069	.1173	.085
.1122	.149	.1193	.050
.1150	.098	.1207	.052
.1220	.117	.1235	.075
.1268	.068	.1256	.079
.1337	.000	.1277	.081
2489180.1007	.121	.1291	.098
.1012	.133	.1304	.102
.1076	.105	.1318	.118
.1101	.118	.1339	.109
.1201	.079	.1353	.152
.1215	.068	.1367	.157
.1285	.012	.1402	.202
.1305	-.118	.1415	.213
.1333	.016	.1429	.252
.1351	.003	.1450	.261
.1360	.007	.1461	.285
.1403	.031	.1470	.286
.1416	.084	.1490	.347
.1465	.013	.1513	.353
.1486	.001	.1526	.378
.1500	.011	.1575	.433
.1514	.031	.1589	.517
.1531	.039	.1603	.508
.1562	.125	.1624	.561
.1601	.206	.1638	.564
.1610	.193	.1651	.577
.1623	.159	.1665	.599
.1615	.110	.1686	.625
.1630	.100	.1719	.676
.1638	.130	.1760	.610
.1701	.132	.1783	.611
.1737	.151	.1818	.611
.1777	.191	.1892	.594
.1826	.242	.1880	.169
.1880	.333	.1901	.453
.1916	.466	.1922	.104
.1986	.589	.1936	.379
.1999	.575	.1985	.327
.2062	.683	.2003	.202
2489181.0770	.081	.2047	.237
.0832	.011	.2061	.189
		.2082	.174
		.2137	.134

TABLE III (Contd.)

J.D.	Δm	J.D.	Δm		
2130107	0811	-227	2139187	1368	.577
	.0889	.240		.1382	.570
	.0917	.204		.1132	.456
	.0952	.113		.1519	.919
	.0991	.500		.1356	.318
	.1008	.326		.1618	.223
	.1077	.629		.1632	.210
	.1105	.670		.1685	.150
	.1119	.091		.1708	.111
	.1133	.753		.1778	.125
	.1160	.738		.1792	.090
	.1174	.762		.1868	.034
	.1188	.790		.1882	-0.010
	.1202	.740		.1938	-0.010
	.1223	.773		.1951	-0.011
	.1237	.751		.1986	.033
	.1250	.769		.2000	.026
	.1271	.743		.2014	.037
	.1290	.727			

Yellow

2439124	1001	.380	2139124	2302	.101
	.1022	.100		.2403	.096
	.1050	.130		.2424	.090
	.1126	.425	2439162	0919	-0.141
	.1151	.402		.0933	.073
	.1181	.351		.0961	.126
	.1258	.231		.1002	.132
	.1292	.157		.1009	.171
	.1320	.117		.1099	.182
	.1383	-0.011		.1120	.186
	.1410	-0.026		.1131	.177
	.1430	.012		.1176	.204
	.1605	.195		.1190	.268
	.1626	.203		.1201	.203
	.1646	.225		.1238	.220
	.1711	.108		.1259	.227
	.1764	.118		.1266	.271
	.1869	.250		.1315	.211
	.1903	.247		.1312	.191
	.1931	.233		.1410	.265
	.1945	.253		.1446	.266
	.2021	.210		.1453	.282
	.2012	.240		.1474	.141
	.2001	.252		.1600	.088
	.2111	.231		.1627	.131
	.2160	.212		.1710	.073
	.2202	.202		.1731	.043

TABLE III (Contd.)

J.D.	Δm	J.D.	Δm
2199160-1990	-1.017	2199167-2129	-.006
.2026	.119	.2156	.041
.2039	.129	.2212	.119
.2075	.121	.2226	.175
.2096	.185	.2271	.171
.2116	.197	.2289	.190
.2130	.269	2199171-1019	+ 2.70
.2151	.330	.1018	.214
.2163	.515	.1981	.119
.2179	.462	.1413	-.116
.2193	.489	.1161	.141
.2207	.502	1569	.319
.2255	.458	1582	.270
.2269	.172	.1906	.228
.2283	.303	.1651	.205
.2297	.378	.1650	.288
.2325	.386	1707	.273
.2339	.365	.1721	.200
.2340	.233	.1763	.211
.2408	.266	.1789	.220
2199167-1510	-.001	.1092	.212
.1523	.062	.1859	.246
.1560	.032	.1908	.236
.1573	+ .008	.1990	.241
.1591	.039	.1971	.240
.1608	.089	.2026	.237
.1650	.191	.2031	.216
.1637	.143	.2068	.226
.1671	.156	.2130	.210
.1691	.207	.2159	.197
.1703	.241	.2172	.191
.1712	.268	.2231	.166
.1726	.275	2235	.114
.1754	.319	.2311	.051
.1768	.371	.2332	-.039
.1702	.303	.2332	+ .022
.1796	.426	.2373	.036
.1809	.406	.2387	.097
.1830	.417	.2408	.141
.1814	.412	.2450	.190
.1838	.433	2199175-0852	.153
.1900	.142	.0887	.096
.1920	.407	.0981	-.116
.1931	.381	.1019	.142
.1918	.397	.1414	.307
.1962	.317	.1135	.290
.1976	.311	.1484	.192
.2025	.222	.1592	.217
.2038	.186	.1615	.248
.2043	.147	.1629	.221
.2059	.139	.1727	.228
.2079	.120	.1747	.168
.2108	.028	.1865	.117

TABLE III (Contd.)

J.D.	Δ^m	J.D.	Δ^m
2139175-1806	-097	2139180-1781	-112
.1007	-093	.1893	-079
.1928	-098	.1995	+013
.1949	.111	.1923	077
.1969	-044	.1992	.211
.1997	+010	.2013	212
.2018	.018	.2069	319
.2016	-100		
.2115	.242	2439181-0742	-176
.2196	.321	.0763	.191
		.0784	.209
2439176-0794	-203	.0846	.258
.0782	.100	.0860	.211
.0844	.248	.0923	.291
.0858	.231	.0944	.291
.0886	.218	.1013	.220
.0907	.214	.1027	.236
.0954	.272	.1096	.239
.0969	.281	.1124	.229
.1010	.317	.1138	.229
.1015	.387	.1159	.211
.1129	.193	.1166	.215
.1157	.226	.1200	.207
.1210	.197	.1221	.208
.1275	.291	.1248	.202
.1311	.273	.1270	185
2439180-1014	.176	.1291	.177
.1010	.174	.1304	.170
.1089	.134	.1318	.168
.1111	.214	.1332	.168
.1187	.218	.1353	.140
.1208	.237	.1367	.111
.1222	.251	.1381	.116
.1282	.334	.1415	077
.1312	.307	.1429	.064
.1340	.311	.1443	036
.1361	.309	.1464	.032
.1375	.314	.1478	.028
.1410	.273	.1492	-007
.1429	.240	.1513	+007
.1472	.282	.1527	.033
.1486	.252	.1540	.094
.1507	.290	.1589	.173
.1521	.274	.1603	.181
.1541	.279	.1617	.227
.1569	.203	.1638	.230
.1611	.092	.1651	.259
.1625	.150	.1665	.270
.1639	.207	.1678	.288
.1632	.220	.1706	.330
.1673	.193	.1769	.366
.1687	.194	.1783	.340
.1708	.174	.1797	.332
.1763	.158	.1832	.303

TABLE III (Contd.)

J. D.	Δm	J. D.	Δm
2199181-1846	.261	2199187-1290	-.499
.1901	-.128	.1214	-.412
.1936	.096	.1250	-.442
.1998	-.016	.1278	-.441
.2019	.043	.1299	.383
.2061	.069	.1375	-.290
.2075	.091	.1389	-.222
.2096	.106	.1459	+ .102
.2131	.157	.1536	-.025
2199187-0018	.131	.1563	-.026
.0030	.105	.1625	.072
.0924	-.007	.1639	.093
.0959	+ .075	.1702	.122
.1001	.153	.1715	.138
.1013	.172	.1785	.173
.1084	.324	.1799	-.202
.1112	.365	.1875	-.260
.1126	.376	.1889	.266
.1153	.419	.1911	-.301
.1167	.433	.1951	.278
.1181	.437	.1993	.259
.1195	.426	.2007	-.213
.1209	.457	.2021	-.257

TABLE V

Phase	m^h	n	Phase	m^h	n
.019	776	10	.013	1420	7
.099	700	6	.026	1371	7
.055	510	6	.017	1251	6
.115	283	1	.066	1112	5
.175	106	11	.110	-0.52	5
.203	027	10	.175	-102	11
.305	016	15	.233	-266	10
.360	110	11	.306	-231	21
.103	211	9	.370	-161	12
.127	350	5	.400	-012	9
.415	157	1	.110	101	12
.160	550	11	.165	1231	10
.103	020	8	.482	1317	1
.500	662	3	.503	1371	7
.516	625	5	.512	1327	1
.350	110	4	.510	1221	5
.300	205	3	.563	1001	5
.611	149	5	.391	1011	7
.703	023	6	.617	136	0
.777	030	8	.605	1230	17
.011	007	7	.770	-219	16
.000	250	7	.037	207	13
.025	133	1	.093	-070	9
.093	520	8	.092	1162	12
.062	651	3	.073	-300	14
.005	766	0	.097	1461	6

TABLE IV

Phase	m _b	m _v	m _b - m _v
.000 m ₁	-1 0m. 790	1 0m. 465	1 0 325
.250	+ 0.005	-0.270	1 0.275
.500 m ₂	+0.662	1 0.975	1 0.287
.730	+ 0.005	- 0 262	1 0 267

To get an idea of the magnitude of uncertainty in the determination of our period we may note that the periods have been calculated over 20,000 cycles. An error of one minute in the minima determination which is expected in a photoelectric method will give an uncertainty of 10^d . 00000008. The earlier photoelectric observers also have indicated comparable figures. The photographic observations by Jensch and Bodokia were somewhat handicapped by lack of well determined minima timings over a long interval and the uncertainties may be estimated to about 10^d . 0000012 on similar basis.

In Table III, the observed differences in magnitudes in two colours between the variable and the comparison stars have been tabulated against the heliocentric Julian moments.

Normal points of the light curve were calculated from 173 comparisons in blue light and 253 comparisons in yellow light from data recorded in Table IV. Figure 2 shows the normal light curve in blue and yellow light respectively.

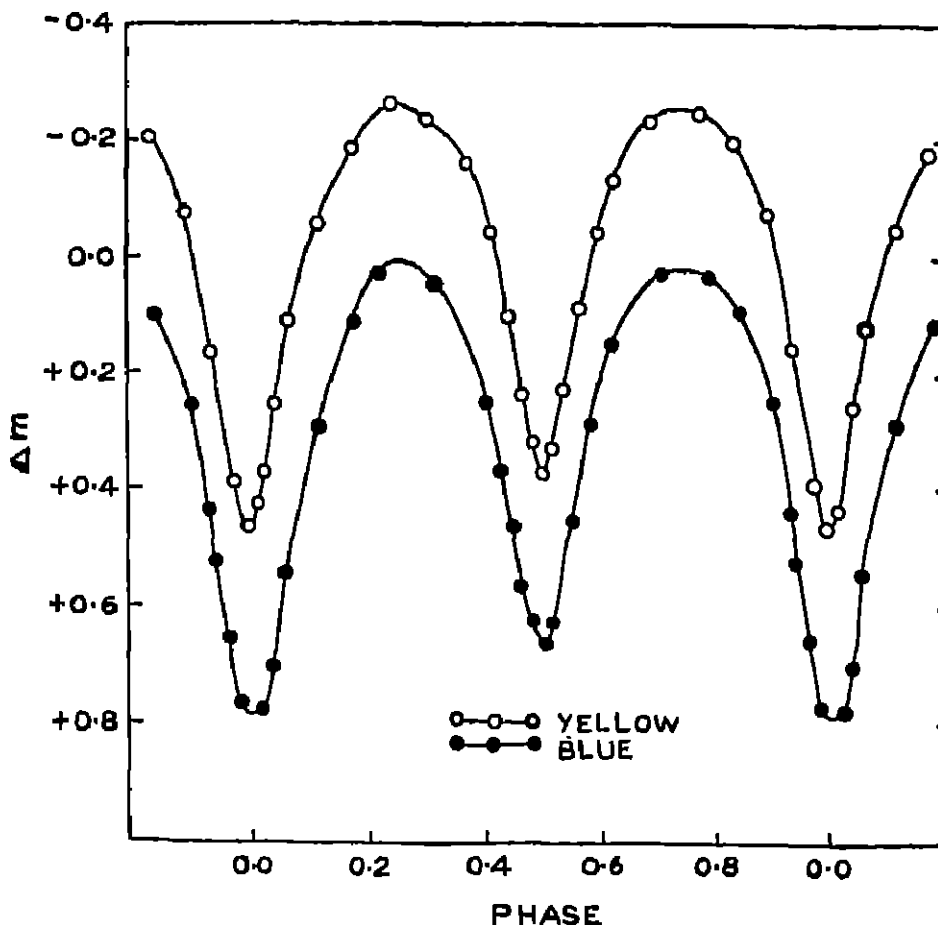


Fig. 2—Light Curve of YY Eridani

From the normal curve the maxima and minima values have been determined and tabulated in Table V. It is seen that the amplitude of the primary minimum exceeds that of the secondary minimum by $+0^m.13$ in blue and $+0.20$ in yellow while the brightness of the system is the same outside eclipses at either elongation.

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