

OBSERVATIONS OF THE BRIGHTNESS TEMPERATURE DISTRIBUTION OF  
THE QUIET SOLAR CORONA AT DECA-METRIC WAVELENGTHS

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The brightness temperature distribution of the quiet solar corona at a wavelength of 8.9 Meters, measured with two types of radio telescopes (1) A "T" type array with a resolution of  $26' \times 38'$  and (2) A fan beam interferometer with an E-W resolution of  $3''$ , found that the persistent bright region do not have any angular structure on scales of  $6''$  or less. The daily variations of the brightness temperature of different regions are studied and the possible interpretation is discussed.