## POSSIBILITY OF OCCURRENCE OF SOLAR PROTON FLARES IN RELATION TO OTHER PHENOMENA

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In the present investigation we have studied the relation between solar proton flares and as ociated optical, radio and X ray events. It is found that optical flares are usually associated with large optical flares in H alpha ( $\geq$  2B) and long duration type II and IV, microwave and hard X ray bursts. The results obtained in the present analy is may be u ed for the prediction of solar proton flare.

During January 1976 May 1986 GOES satellite recorded 60 proton (> 10 MeV) flare events. Out of these 8 events were excluded from our analysis because no flares have been recorded in H alpha observations. We found that 77% proton flares are associated with H alpha flares of importance class > 28. We have calculated flare index (Kieczek 1952) for various proton flare producing active regions. We found that proton flare start at that day when flare index is maximum in most of the active regions. We have studied association between proton flares and various types of radio bursts. It is found that about 70% type II, III and IV radio bursts are associated with proton flares. Type I and V radio bursts show poor association (38% and 29%) with proton flares. The duration of radio bursts is generally large, type I > 30 minutes, type II > 10 minutes, type III & IV > 5 minutes type IV > 30 minutes. The proton flares show 6.7% association with the microwave bursts (range 0.9.2.0, 2.9.5.0 and 8.0.11.0 GHz. The durations of proton flare associated microwave bursts are greater than 20 minutes. We have found that 61% proton flares show association with hard X ray bursts (> 25 KeV). The duration of proton flares associated. HXR bursts are > 6 min.

The results obtained here may be used for the prediction of proton flares since H alpha flares, microwave, radio bursts and HXR bursts are observed much before (~8 hrs) the proton flares. Therefore proton flares may be predicted upto 8 hrs before as observed by satellite.

The detailed analysis of this work is under publication elsewhere

## Reference

Kleczek J 1952 Bull Astron Inst Czech 3, 52