## THE COMET OF NOVEMBER, 1948*

THE recent comet which according to newspaper reports was first sighted by the pilot of an Australian air liner on the morning of 7 -II-48 is one of the brightest that have been seen since the last appearance of Halley's comet in IgIO. At Kodaikanal, on account of unfavourable weather conditions, observations of the comet could be begun only on ri-mr-48, but they were continued till I4-T2-48, with several interruptions caused by bad weather. By the 14th, the comet had become invisible to the naked eye, and telescopic and photographic observations were also not possible subsequently because of bad weather.

In all, visual and photographic observations of the comet were made on I3 days, while on two other days only visual observations were possible. The results obtained are summarised below. Table I gives the approximate right ascension and declina tion of the nucleus of the comet as computed from the photographs with reference to the surrounding star field.

From the table it will be seen that the total movement in right ascension from $1 \mathrm{I}-\mathrm{II}-48$ to

[^0]7-12-48 was $3 \mathrm{~h} .36 \mathrm{~m} .\left(=54^{\circ}\right)$, and that in declination $\mathrm{Ir}^{\circ} 34^{\prime}$. If the observed coordinates are plotted

TABLE I

| Date. | Mean time of photograph (hrs. IST). | Right Ascension of comet. | Declination of comet. |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | h. m. | h. m. |  |  |  |
| 11-11-48 | 0530 | 1303 | $23^{\circ}$ | $24^{\prime}$ | (South) |
| 12-11-48 | 0500 | $12 \quad 54$ | $24^{\circ}$ | $18^{\prime}$ |  |
| 13-11-48 | 0500 | 1247 | $25^{\circ}$ | $08^{\prime}$ | " |
| 19-11-48 | 0430 | 1158 | $29^{\circ}$ | $20^{\prime}$ | " |
| 21-11-48 | 0430 | 1140 | $30^{\circ}$ | $32^{\prime}$ | " |
| 24-1 1-48 | 0500 | 1116 | $32^{\circ}$ | $00^{\prime}$ | " |
| 25-11-48 | 0430 | 1108 | $32^{\circ}$ | $30^{\prime}$ | " |
| 26-11-48 | 0445 | $10 \quad 59$ | $32^{\circ}$ | $56^{\prime}$ | " |
| 27-11-48 | 0500 | 1051 | $33^{\circ}$ | $17^{\prime}$ | " |
| 28-1 1-48 | 0330 | 1043 | $33^{\circ}$ | $37^{\prime}$ | n |
| 29-11-48 | $64 \quad 00$ | $10 \quad 34$ | $33^{\circ}$ | $58^{\prime}$ | " |
| 1-12-48 | $04 \quad 00$ | $10 \quad 17$ | $34^{\circ}$ | $30^{\prime}$ | " |
| 7-12-48 | 0245 | 0927 | .$^{34}$ |  | " |

against the corresponding dates (vide Fig. I) and smoothed curves drawn, it will be seen that whilile
the rate of movement in right ascension was quite uniform that in declination steadily decreased during the period. The rate of movement in right ascension obtained from Curve I, Fig. 1 is 8.3 min . per day $\left(2^{\circ} 4^{\prime \prime} 5\right)$. Fig. 2 gives a graphical representation of the daily rates of movement in declination obtained by plotting the values taken from Curve II, Fig. I against the corresponding dates. The rates taken from the smoothed curve in Fig. 2 are given in Table II.

It will be seen that the rate of movement in declination decreased from nearly a degree per day at the beginning to less than a minute towards the end of the period. Fig. 3 shows the track of the comet against the background of fixed stars. Typical photographs of the comet are reproduced in Fig. 4.

The tail of the comet when first sighted subtended an angle of $28^{\circ}$ and was pointing approxi-

TABLE II

| Date interval | Rate of movement in declination | Date interval | Rate of movement in declination |
| :---: | :---: | :---: | :---: |
| 11-12 Nov. 48 | 54 | 24-25 Nov. 48 | 27 |
| 12-13 " | 51 | 25-26 , | 25 |
| 13-14 , | $48 \cdot 3$ | 26-27 ", | 23 |
| 14-15 ", | $45 \cdot 6$ | 27-28 ", | 21 |
| 15-16 ", | 43 | 28-29 , | $18 \cdot 8$ |
| 16-17 , | 41 | 29-30 ", | 16 |
| 17-18 ", | 39 | 30 Nov.-1 Dec. 48 | 13 |
| 18-19 ", | 37 | 1-2 Dec. 48 | 10 |
| 19-20 ", | $35 \cdot 3$ | 2-3 , | $7 \cdot 5$ |
| 20-21 ", | $33 \cdot 7$ | 3-4 " | $5 \cdot 3$ |
| 21-22 ", | 32 | 4-5 ," | 3.4 |
| 22-23 ", | $30 \cdot 4$ | 5-6 ," | 1.8 |
| 23-24 " | $28 \cdot 6$ | 6-7 \% | $0 \cdot 5$ |

mately in a soutbwesterly direction. The brightness of the head at that time was estimated at about $I^{\circ} 5$


FIC. 1
stellar magnitude. On 7-12-48 when the comet was last photographed it had become very faint (about magnitude 5) and the tail had dwindled to less than $5^{\circ}$. The computed angular distance of the comet from Venus on II-II-48 was $20^{\circ} 48^{\prime}$, the correspond.
ing figure on 7 -12-48 was $73^{\circ} 3^{6^{\prime}}$. The direction of movement of the comet was approximately westsouthwestwards. On $14-\times 2-48$ when the comet was last seen through the telescope at Kodaikanal it had entered the region of the Milky Way.


Fig. 2




[^0]:    * Communicated by the Solar Physics Observatory, Kodaikanal.

