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SOLUTION OF RADIATIVE TRANSFER EQUATION IN SPHERICALLY SYMMETRIC MEDIA WITH SPHERICAL HARMONIC APPROXIMATION

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ABOTRACT

A numerical method for obtaining solution of radiative transfer equation in spherically symmetric media with spherical harmonic approximation is presented. The angle derivative is approximated by an orthonormal polynomial and this is represented by a matrix called curvature matrix, for a given beam of rays. An error analysis of the curvature matrix and results using the solution for few representative passes have been presented.

Key words : spherical symmetry-radiative transfer equation-spherical harmonic approximation

1. Introduction

Recently, several techniques for getting solution of radiative transfer in spherically symmetric media have appeared in the literature (see Hummer and Rybicki 1971, Peralah and Grant 1973 and others). However, each technique has its own advantage and disadvantage. Moreover, these techniques have to be modified according to the given physical situation. In this respect, the discrete space theory (Peralah and Grant 1973) is one method which needs the least modification. The average specific intensities in the "cell" are approximated by the "diamond" scheme which proved to be the best choice (Wiscombs 1976). The transmission and reflection characteristics of the medium are directly dependent on its physics. These are ironted as operators subjected to numerical analysis. The main condition for these operators is that they should be nonnegative. For this condition to be satisfied in spherical geometry, we have to select a small "cell" critical thickness because of the fact that the two curvature matrices are asymmetric and large in their absolute values. To avoid this difficulty, we have tried to obtain the curvature matrix by the spherical harmonic method.

2. Development of the Method

2.1 Derivation of Transmission and Reflection Operators

The equation of radiative transfer in apherical symmetry is written as,

$$\frac{\mu}{r^{2}} \frac{\partial}{\partial r} \left\{ r^{2} \mathbf{I} (r, \mu) \right\} + \frac{1}{r} \frac{\partial}{\partial \mu} \left\{ (1 - \mu^{2}) \mathbf{I} (r, \mu) \right\} + K (r) \mathbf{I} (r, \mu) = K (r) \left\{ \mathbf{I} - \omega(r) \right\} \mathbf{B} (r) + \frac{1}{r} \omega (r) \int_{-1}^{+1} \mathbf{P} (r, \mu, \mu') \mathbf{I} (r, \mu') d\mu' \right\}$$
(1)

where K (r) is the absorption coefficient and K (r) ≈ 0 and the ω (r) is the albedo for single scattering $0 \leq \omega$ (r) ≤ 1 . I (r, μ) is the monochromatic specific intensity of the ray making an angle $\cos^{-1} \mu$ with the radius vector at the radial point r. B (r) is the Planck function at r and P (r, μ , μ') is the phase function and we shall choose phase function for isotropic scattering in these calculations. The angles are discretized such that $0 < \mu$, $< \mu_2 < ... < \mu_m \leq 1$. We shall write,

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$$I_{j}^{*}(r) = I(r, \mu_{j}), I^{*}_{j}(r) = I(r, -\mu_{j})$$
 (2)

and

$$\mathbf{M} = [\mu_1 \ \delta_{13}] \tag{3}$$

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We have to approximate $\frac{\delta I}{\delta \mu}$ within this discretization of the angular coordinate μ . Let $P_m(\mu)$, μ_{j} , C_l be orthonormal polynomials, abscisses and weights associated with Mth order Gauss rule on [0, 1]. We shall expand I (μ) as,

$$I(\mu) = \sum_{m=0}^{M} a_m P_m(\mu)$$
 (4)

and the (m+1) coefficients are chosen so that

$$T(\mu_j) = \sum_{m=0}^{M} a_m P_m(\mu_j), j = 0, 1, ..., M$$
 (6)

Since, the P_m (μ) are assumed orthonormal

$$\sum_{j=0}^{M} P_{m}(\mu_{j}) P_{m}(\mu_{j}) C_{j} = \delta_{k,m}$$

$$(6)$$

From equation (6), we have

$$a_{\rm m} = \sum_{j=0}^{\rm M} P_{\rm m} (\mu_{\rm j}) C_{\rm j}, I(\mu_{\rm j})$$
 (7)

(see Abramowitz and Stegun, 1964).

From equation (4), we have,

$$\begin{array}{c} 1'(\mu) \approx \sum\limits_{m=0}^{M} a_m P_m'(\mu) \end{array}$$

where the primes represent the derivatives with respect to μ . Therefore from equations (5) and (7), we obtain,

1'
$$(\mu_j) \approx \frac{M}{\Sigma} \sum_{\substack{j'=0 \\ m \neq 0}}^{M-1} Pm' (\mu_j) Cj, Pm (\mu_{j,j}) T (\mu_{j,j})$$

or

$$\frac{dI}{d\mu} = D I$$

$$Djj, = \sum_{m=0}^{M} Pm'(\mu j) Cj, Pm(\mu j.)$$
(8)

Consequently,

$$-\frac{\hbar}{2}\frac{\partial}{\partial\mu}\left\{(1-\mu^{2})\mathbf{I}^{*}(r,\mu)\right\} \left| \mu-\mu_{1}-\mu_{1}\mathbf{I}^{*}(r,\mu) -\frac{1}{4}(1-\mu)^{2}\right)\frac{d\mathbf{I}^{*}(r,\mu)}{d\mu}$$
(9)

We shall introduce a matrix operator A called the "curvature scattering matrix" given by

$$\Lambda = [\Lambda_{\rm He}] \tag{10}$$

8nd

$$\Lambda_{II} = \mu_{I} \, \delta_{II}, \, - \frac{1}{4} \, (I - \mu_{I}^{4}) \, \mathsf{D}_{II}, \tag{11}$$

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Now, equation (9) can be rewritten as,

$$-\frac{1}{2} \frac{\partial}{\partial \mu} \left\{ (1 - \mu^2) I^*(\mathbf{r}, \mu) \right\} \bigg|_{\mu = \mu_1} - \sum \Lambda_{ij} I^*(\mathbf{r}, \mu)$$
(12)

or alternetively,

$$-\frac{1}{2}\frac{\partial}{\partial\mu}\left\{\begin{array}{c}(1-\mu^2)\ 1^+(r,\mu)\end{array}\right\}\approx \sum_{m=0}^{M}a_{n}\left\{-\frac{1}{2}\ \frac{\partial}{\partial\mu}\left[(1-\mu^2)\ P_{m}(\mu)\right]\end{array}\right\}$$

so that

$$A_{jj} = \sum_{m=0}^{M} \frac{1}{2} \frac{d}{d\mu} [(1-\mu^2) P_m(\mu)] \Big|_{\mu = \mu_j} C_j, P_m(\mu_j)$$
(13)

For computational purposes, we shall use the expression (11) instead of (13).

The equation of radiative transfer equation (eq 1) for a given beam of radiation is given by

$$M \frac{1}{r^2} \frac{\partial}{\partial r} \left\{ r^2 l^+(r) \right\} - \frac{2}{r} \Lambda l^+(r) + K(r) l^+(r) = K(r) \left\{ \left[1 - \omega(r) \right] B^+(r) + \frac{1}{2} \omega(r) \left[P^{++}(r) C l^+(r) + P^{+-}(r) C^{-}(r) \right] \right\}$$

and for the oppositely directed beam,

$$- M - \frac{1}{r^2} \frac{\partial}{\partial r} \left\{ r^2 I^-(r) \right\} - \frac{2}{r} \wedge I^-(r) + K (r) I^-(r) - K (r) \left\{ \left[1 - \omega (r) \right] B (r) + \frac{1}{2} \omega (r) \left[P^{-+}(r) C I^+(r) + P^-(r) C I^-(r) \right] \right\}$$
(14)

We shall define

$$U (r) - A (r) I (r)$$

$$A (r) - 4\pi r^{2},$$

$$b \pm (r) - A (r) B \pm (r)$$
(15)

By introducing (15) Into (14), we obtain,

$$M \frac{dU(r)}{dr} + K(r) U^{+}(r) - K(r) [1 - \omega(r)] b^{+}(r) + [\frac{1}{4} K(r) \omega(r) P^{++}C + \frac{2}{r} \Lambda] U^{+}(r) + \frac{1}{4} K(r) \omega(r) P^{+-}C U^{-}(r)$$

and

$$-M \frac{d U^{-}(r)}{dr} + K (r) U^{-} (r) - K (r) [1 - \omega (r)] b^{+} (r) + \frac{1}{4} K (r) \omega (r) P^{-+} C U^{+} (r) + [\frac{1}{4} K (r) \omega (r) P^{--} C - \frac{2}{r} \Lambda] U^{-} (r)$$
(16)

Discretization in the radial coordinate gives us,

$$\mathbf{M} \left(\mathbf{U}_{i}^{-} - \mathbf{U}_{i+1}^{-} \right] + \tau_{i+1} \left[\mathbf{U}_{i+1}^{-} - \tau_{i+1} \left[\left(1 - \omega_{i+1} \right) \mathbf{b}_{i+1}^{-} + \mathbf{Q}_{i+1}^{-+} \mathbf{U}_{i+1}^{+} + \mathbf{Q}_{i-1+1}^{--} \mathbf{U}_{i+1}^{--} \right] \mathbf{U}_{i+1}^{--}$$
are

$$\tau_{i+1} = K_{i+1} (r_{i+1} - r_i)$$
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$$Q^{++}_{i+1} = \frac{1}{2} \omega_{i+1} P^{++}_{i+1} C + 2 \frac{\varphi}{\tau_{i+1}} \Lambda$$

$$Q^{--}_{i+1} = \frac{1}{2} \omega_{i+1} P^{--}_{i+1} C - 2 \frac{\varphi}{\tau_{i+1}} \Lambda$$

$$Q^{+-}_{i+1} = \frac{1}{2} \omega_{i+1} P^{+-}_{i+1} C$$

$$Q^{-+}_{i+1} = \frac{1}{2} \omega_{i+1} P^{-+}_{i+1} C$$

$$U^{\dagger}_{i+1} = \frac{1}{2} (U^{\dagger}_{i+} + U^{\dagger}_{i})$$
(20)

where g is the curvature factor given by

Substituting the last equation (20) in (17) and (18) and then writing these difference equations in the form of the interaction principle (see Peraleh and Grant 1973) would give us,

$$\begin{bmatrix} M + \frac{1}{2} \tau_{i+1} (I - Q^{i+1}_{i+1}) & -\frac{1}{2} \tau_{i+1} Q^{i-1}_{i+1} \\ -\frac{1}{2} \tau_{i+1} Q^{-1}_{i+1} & M + \frac{1}{2} \tau_{i+1} (I - Q^{-1}_{i+1}) \end{bmatrix} \begin{bmatrix} U^{i}_{i+1} \\ U^{i}_{i} \end{bmatrix} + \tau_{i+1} (1 - \omega_{i+1}) \begin{bmatrix} b^{i}_{i+1} \\ b_{i+1} \end{bmatrix} + \begin{bmatrix} M - \frac{1}{2} \tau_{i+1} (I - Q^{i+1}_{i+1}) & \frac{1}{2} \tau_{i+1} Q^{i-1}_{i+1} \\ \frac{1}{2} \tau_{i+1} Q^{-1}_{i+1} & M - \frac{1}{2} \tau_{i+1} (I - Q^{-1}_{i+1}) \end{bmatrix} \begin{bmatrix} U_{i}^{i} \\ U^{i}_{i+1} \end{bmatrix}$$
(21)

Here I is the identity matrix.

Therefore, the output intensities can be rewritten in terms of the input intensities as,

$$\begin{bmatrix} U^{+}_{i+1} \\ U^{-}_{i} \end{bmatrix} \sim K^{-1} \tau_{i+1} (1 - \omega_{i+1}) \begin{bmatrix} b^{+}_{i+1} \\ b^{-}_{i+1} \end{bmatrix} + K^{-1} \begin{bmatrix} M - \frac{1}{2} \tau_{i+1} (1 - Q^{++}_{i+1}) & \frac{1}{2} \tau_{i+1} Q^{+-}_{i+1} \\ \frac{1}{2} \tau_{i+1} Q^{-+}_{i+1} & M - \frac{1}{2} \tau_{i+1} (1 - Q^{--}_{i+1}) \end{bmatrix} \begin{bmatrix} U^{+}_{i} \\ U^{-}_{i+1} \end{bmatrix}$$
(22)

where,

$$K = \begin{bmatrix} M - \frac{1}{4} \tau_{i+\frac{1}{4}} (I - Q^{++}_{i+\frac{1}{4}}) & -\frac{1}{4} \tau_{i+\frac{1}{4}} Q^{+-}_{i+\frac{1}{4}} \\ -\frac{1}{4} \tau_{i+\frac{1}{4}} Q^{-+}_{i+\frac{1}{4}} & M + \frac{1}{4} \tau_{i+\frac{1}{4}} (I - Q^{--}_{i+\frac{1}{4}}) \end{bmatrix}$$
(23)

and

$$\mathbf{K}^{-1} = \begin{bmatrix} (\mathbf{I} - \mathbf{r}^{+-} \mathbf{r}^{++})^{-1} \Delta^{+} & (\mathbf{I} - \mathbf{r}^{+-} \mathbf{r}^{-+})^{-1} \mathbf{r}^{+-} \Delta^{-} \\ (\mathbf{I} - \mathbf{r}^{-+} \mathbf{r}^{+-})^{-1} \mathbf{r}^{-+} \Delta^{+} & (\mathbf{I} - \mathbf{r}^{-+} \mathbf{r}^{+-})^{-1} \Delta^{-} \end{bmatrix}$$
(24)

where

$$\Delta^{+} = [\mathbf{M} + \frac{1}{2} \tau_{i_{+}i_{+}} (I - \Omega^{++}_{i_{+}i_{+}})]^{-1}$$

$$\Delta^{-} = [\mathbf{M} + \frac{1}{2} \tau_{i_{+}i_{+}} (I - \Omega^{--}_{i_{+}i_{+}})]^{-1}$$

$$r^{+-} = \frac{1}{2} \tau_{i_{+}i_{+}} \Delta^{+} \Omega^{+-}_{i_{+}i_{+}i_{+}}$$

$$r^{-+} = \frac{1}{2} \tau_{i_{+}i_{+}} \Delta^{-} \Omega^{-+}_{i_{+}i_{+}i_{+}i_{+}}$$
(25)

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compare equation (22) with the local Interaction principle

 $\begin{bmatrix} \mathbf{U}^{+}_{i+1} \\ \mathbf{U}^{-}_{i} \end{bmatrix} = \begin{bmatrix} \mathbf{t} (i+1,i) & \mathbf{r} (i,i+1) \\ \mathbf{r} (i+1,i) & \mathbf{t} (i,i+1) \end{bmatrix} \begin{bmatrix} \mathbf{U}^{+}_{i} \\ \mathbf{U}^{-}_{i+1} \end{bmatrix} + \begin{bmatrix} \boldsymbol{\Sigma}^{+}_{i+1} \\ \boldsymbol{\Sigma}^{-}_{i+1} \end{bmatrix}$

and setting

 $A^{++} = M - \frac{1}{2} \tau_{l+1} (I - Q^{++}_{l+1})$ $A^{--} = M - \frac{1}{2} \tau_{i+1} (I - Q^{--}_{i+1})$

We obtain the transmission and reflection operators in the basic "cell" and they are given by,

t (l+1, l) - R⁺⁻ [Δ⁺ A⁺⁺ + r⁺⁻ Δ⁻ B⁻⁺]

B"* - 1 + Q"*I+1

t (l, i+1) - R⁻⁺ [Δ⁻ A⁻⁻ + r⁻⁺ Δ⁺ B⁺⁻] $r (I+1, I) = R^{-+} [\Delta^{-} B^{-+} + r^{-+} \Delta^{+} A^{++}]$ (28) $r(l, l+1) = R^{+-} [\Delta^+ B^{+-} + r^{+-} \Delta^- A^{--}]$

(26)

(27)

And the source vectors are given by,

$$\Sigma^{+}_{1} + \frac{1}{1} = \tau_{1} + \frac{1}{1} (1 - \omega_{1} + \frac{1}{2}) R^{+-} [\Delta^{+} b^{+} + r^{+-} \Delta^{-} b^{-}]$$

$$\Sigma^{-}_{1} + \frac{1}{2} = \tau_{1} + \frac{1}{2} (1 - \omega_{1} + \frac{1}{2}) R^{-+} [\Delta^{-} b^{-} + r^{-+} \Delta^{+} b^{+}]$$
(29)

where

$$\Sigma^{-}(+_{i} = \tau_{i} +_{i} (1 - \omega_{i} +_{i}) R^{-+} [\Delta^{-} b^{-} + r^{-+} \Delta^{+} b^{+}]$$

$$R^{+-} = [L - r^{+-} r^{-+}]^{-1}$$
(29)

$$\mathbf{R}^{++} = [\mathbf{I} - \mathbf{r}^{-+} \mathbf{r}^{+-}]^{-1}$$
 (30)

Equations (28) to (30) will give us the operators in a "cell".

2.2 Accuracy of the Orthonormal Approximation

In the previous sub-section, we have approximated,

Μ Σ J•

Let

 $f(\mu_{i}) = -\frac{1}{2}\frac{\partial}{\partial\mu} \{(1-\mu^{2}) I(\mu)\}$ (32) $\mu = \mu_{\rm I}$

So, we must calculate the error E involved

$$\mathbf{E} = \begin{bmatrix} \mathbf{f} (\boldsymbol{\mu}_{j}) & -\boldsymbol{\Sigma} \wedge_{jj}, \mathbf{I} (\boldsymbol{\mu}_{j}) \end{bmatrix}$$
(33)

Let us make use of a simple expansion for the specific intensity as some power of μ given by,

I
$$(\mu) = \mu^n$$
, n=0, 1,2,..., N (34)

$$\frac{\partial}{\partial \mu} \{ (1 - \mu^2) | \pm (r, \mu) \} \\ \mu - \mu_1 \\ \mu - \mu_1$$

Substituting (34) into (32) and (33) we obtain,

$$E = \frac{(n+2)\mu_1}{2} - \frac{n\mu_1}{2} - \frac{\Sigma}{L} \wedge \mu_1, \qquad (36)$$

The A_{jj} , are evaluated by using the zeros and weights of Gauss-Legendre quadrature formula over (0,1) for $\mu^{\prime *}$ and C'*. We have considered j, j'=1,...J, J=8. The orthonormal shifted Legendre polynomials are calculated by the recurrence relation

$$P_{m-1}(\mu_{i}) = \frac{2m-3}{m-1} (2\mu_{i}-1) P_{m-2}(\mu_{i}) - \frac{m-2}{m-1} P_{m-3}(\mu_{i}), \quad m = 3, 4, ..., M$$
(36)

with

$$P_0(\mu_j) = 1$$
, and $P_1(\mu_j) = (2\mu_j - 1)$

The orthonormal condition requires that

$$P_{m_{-1}} \mu_{j} = \sqrt{2m-1} P_{m_{-1}} (\mu_{j})$$
(37)

One can easily write the derivatives $\dot{P_m} \sim I(\mu)$ from the relation (36) and (37) The polynomials and their derivatives for J 8 are given in tables 1 and 2. The Λ matrices have been calculated for 4 and 6 angles. These are given in Tables 3 and 6. By making use of these, we evaluate E in equation (35) and is given by

$$E = \begin{cases} n \\ f(\mu_j) - \sum_{j, k \in J} A_{ij}, \mu_j, \\ l, \end{cases} < 10^{-6} \text{ for } n < M - 1$$

and for $n \ge M$, E increases gradually. For n > 8, we have plotted the error in figure 1. For n < 8, the results seem to be almost exact.



Fig. 1 The errors defined in equation (35) are plotted against n. For n = 1 to 7, E=0 (10⁻⁴).

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Table 1. Orthonormal Polynomials for J~8

m	J	1	2	3	4	5	6	7	6
0		1.0000	1.0000	1.0000	1.0000	1.0000	1.0000	1.0000	1.0000
1		-1.6632	-1.3798		3177	.3177	.8102	1.3798	1.6632
2		1.0749	1.0107	1918	-1.0051	-1.0051	— .1018	1.0107	1.9749
3		-2.0482	1827	1.1258	.6871	— 0 871	1,1256	.1827	2.0462
4		1.9118	7281	9809	.7813	.7613	·9809	7281	1.0118
6		-1.8014	1.3376	1028		.9570	.1020	-1.3376	1.8014
8		1.1801	-1.3948	1.0902	4088	.4089	1.0902	- 1.3846	1,1504
7		.6003	.8775		1.1178	-1.1176	1 0388	8775	.5003

Table 2. Derivatives of the orthonormal polynomials given in Table 1

m]	1	2	3	4	8	6	7	8
0		0	0	0	0	0	0	0	0
í		3.4641	3.4641	3 48 41	3 4641	3.484 1	3.4641	3.4841	3.4841
2		-12.8838	10.6884	- 7.0507	2.4810	2.4610	7.0607	10.6884	12.6836
3		28.8587	17.2507	3.0234	-6 6018	-6,6018	3.0234	17.2607	28.6597
4		-49.7884	~17.2407	8 4088	7.8084	-7.6064		17.2407	49.7684
5		73.9728	7.1346	-15 7300	6.8741	6.8741	16.7300	7,1346	73.9728
8		-98.1147	11.2711	7.6463	-13.8859	13.9859	7.6436		BB.1147

Table 3. Curvature costiaring matrices Λ_{11} for $J\!=\!4$

3,3863	-4.8367	2 0775	0.6778	•
0.6750	0.8725	-1.3105	0.2929	
0.1811	0.8108	0.4681	0.4174	
0 0778	0 2826	0.6514	0.4839	

Table 4. Corrected All, for J = 4

2.2880	-4.2519	1.7948	0.4959	-
0.0897	0.9844	-1.4819	0 3385	
0 7864	1,1224	0.3073	0.3738	
L —1 0194	0.3020	0,3687	0.5668	

Table 5. Correction matrix (j), for j=4

1.0972	0.5847	0.2827	0.0918]
0.5652	0.3119	0.1508	0 0438
0.5652	0.3119	0.1508	0.0436
1.0972	0.5847	0.2827	0.0818

Table 6. Metrix for J~8

12.3511 19.6159 1.6837 2.2598 -0.4043 2.0819 0.1608 -0 6542 -0.0652 0.3192 0 0524 0 1891 -0.0342 0.1211 0.0204 0.0718	159 12.3422 → 8 5814 598 6.1022 3.3520 818 0 9224 3.4272 542 1 9808 0.55864 192 0 7378 1.7714 891 0 3978 0.7328 211 0 2440 0 4090 718 0.1420 0.2292	5 8287 3.6144 2.0873 1.2513 1.0637 0.8978 2.2714 0.9480 0.4663 1.5290 1.5187 0.4591 0.6636 1.1900 0.3375 0.4854	1.8288 0.6211 0.4289 0.4083 0.5102 -0.8225 0.4774 0.7715	0.5204 0.1754 0 1183 <i>0.1092</i> 0.1254 0.1791 0 3673 0 4951
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2.3 Conservation of flux and non negativity of transmission and reflection matrices.

Any medium which relther creates nor absorbs radiation, must conserve flux. We shall apply this condition (see Grant and Hunt 1969 a, b) and arrive at the identity that

However, when we apply (38) to the ∧ matrices given in Tables (3) and (6), we find that this identity is not satisfied. We shall, therefore, derive corrections to the curvature matrices so that they always satisfy the flux condition. Let

or

$$\begin{array}{c} J \\ \Sigma \\ C_{j} \left(\Lambda_{jk} - \underline{\epsilon'_{k}} \right) = 0 \\ j = 1 \\ C_{j} \end{array}$$
 (40)

Let

$$\epsilon_{jj} = \frac{\epsilon_{jj}}{C_j}$$

and

$$\Lambda'_{\rm H} = \Lambda_{\rm H}, \quad (41)$$

then, we have

The matrices ι' and Λ' are given in Tables 5 and 8 and 4 and 7 for J = 4 and 8 respectively.

To obtain non-negative transiesion and reflection operators, we must have the critical thickness in the "cell" given by (for positive Δ^+ and Δ^- matrices)

$$\tau_{\rm crit} = \min_{j} \left| \frac{\mu_{j} \pm g \Lambda'_{ij}}{\frac{1}{4} \left(-\frac{1}{4} \omega P^{*+}_{jj} C_{j} \right)} \right|$$
(43)

Results for few simple cases have been presented in the next section.

3. Results

We have used the method to calculate the emergent intensities in two cases. We have used J = 4 and B/A = 1.5, 2 and 2.5 and m = 1 and $\tau = 10$ and 100 the total optical depth, where B and A are the outer and inner radii of the atmosphere. In figures 2 and 3, we have given the angular distribution of the emergent radiation 1 (0, μ) ($\sim U$ (0, U) $/4\pi B^3$). We notice that as B/A increases, the emergent intensities reduce. This is nothing but the dilution due to extendedness and curvature scattering. When the optical depth is increased from 10 to 100, we notice a considerable fall in the emergent intensities which is quite obvious on the physical grounds.



Fig. 2 Angular distribution of the emergent radiation for $\tau \approx 10$.



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Table 7, Corrected 📙 for J 🖙 8

10.3965	- 18.3928	11.4667	-7.9490	5.3920	-3.3387	1.6885	-0.4808
0.9952	2 8165	- 6.5006	3 6398	-2.2960	1.3764	-0.6840	0.1934
-1 0334	2.4767	0.6399	-3.2231	1.5128	0.8098	0.3823	- 0.1066
- 0.3842	0.3128	1.7163	0.7429	- 2.3933	1.0221	-0.4479	0.1202
- 0.6302	0.6606	- 0.9821	1.8479	0.3464	1.4529	0.4716	0.1144
0.6777	0.2067	0.1153	-0.6287	1.3778	0 5471	- 0 9671	0.1919
- 0.9227	0.6778	- 0.6424	0.8988	-0.8623	1.3141	0 4146	- 0.3493
- 1.9322	1 1615	-0.7366	0.4032	-0.0092	0.2127	0.6332	0 6346

Table 8. Correction matrix jj for Jan8

1.9626	-1.2233	0.8765	-0.8324	0.4367	0,2727	0.1383	0 0396
0.8685	0.6567	0 3984	- 0.2878	0.1987	· 0.1241	0.0629	-0.0180
0.6301	0.3946	0.2825	-0.2041	0.1408	- 0.0880	0 0446	0 0128
0.5450	- 0 3414	0.2443	0.1765	0.1219	0 0761	0.0386	0 0110
0.6460	0 3414	0.2443	0,1786	0.1219	0 0761	0.0386	-0.0110
0.6301	-0.3948	0.2826	-0.2041	0.1409	- 0 088 0	0 0448	- 0.0126
0.8886	- 0.6687	0.3984	-0.2878	0.1987	0.1241	0.0829	- 0.0180
1.9620	1 2233	0.8765	-0.8324	0.4307	- 0.2727	0.1383	- 0.0396

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References

Abramowitz, M., Stegun, I.A. 1984, Handbook of Mathematical Functions, Dover, New York, Grant, I.P., Huni, G.E., 1989a, Proc. Roy. Soc. London Ser. A, 313, 183. Grant, I.P., Hunt, G.E., 1989b, Proc. Roy. Soc. London Ser. A, 313, 199 Hummer, D.G., Ryblaki, G.B. 1971, Mon. Not. R. Astr. Soc. 182, 1. Peralah, A., Grant, I.P., 1973 J. Inst. Meth. Appless. 12, 76. Wiscombs. W.J., 1976, JOSRT, 18, 637.