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Fifteen hypervelocity stars (HVS) whose Gaia G magnitudes are less than 16.0 and radial velocities less than -600 km per second

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Abstract

From Gaia DR3 data using a simple search criteria fifteen hypervelocity stars (HVS) are detected. These stars have radial velocities

less than -600 km per second, Gaia G magnitudes less than 16.0 and RUWE values less than 1.4. The U, V, W

velocities also indicate that these stars are HVS.

The effective temperatures of these stars indicate that they are

F and G type stars.

Gaia DR3 409420157955913856 and Gaia DR3 2198292118993038464

seems to be metal-poor HVS. High resolution

spectroscopy of these stars is needed to derive the atmospheric parameters,

chemical composition and to understand their

evolutionary status and origin.

The progenitors of these stars may be close binary systems with white dwarf companions. When the white dwarf companions became Type Ia supernovae and their F-G companion stars have become HVS. Or they may be the result of other unknown processes.

Keywords: Hypervelocity stars (HVS) - Parallaxes - Radial Velocities - Gaia DR3

Introduction

The first hypervelocity star (HVS) was discovered by Brown et al. (2005) (see Hills 1988). Lateron from systematic surveys many HVS were discovered. Our Galactic Center (GC) may produce HVS by various processes. Brown (2015) published a review paper on HVS. The presence of HVS beyond the GC may be high velocity runaways ejected by supernovae in close-binary stars (Blaauw 1961). The causes for the presence of many HVS and high velocity stars in the GC and all over in the Galaxy are still not well understood. Gaia DR3 has revealed many HVS which include HVS white dwarfs (Gaia Collaboration et al. 2021, 2022). (see also Shen et al. 2018, Evans et al. 2023 (references therein), Liao et al. 2023 (refrences therein)). In Refrise paper of the American Astronomical Society using a simple search criteria.

Selected HVS candidates

I have searched the Gaia DR3 data for HVS with radial velocities (RV) less than -600.0 km per second, Gaia G magnitudes less than 16.0 and RUWE (Renormalized Unit Weight Error) values less than 1.4. The RUWE is the primary astrometric goodness-of-fit indicator (Lindegren et al. 2021). The choice of less than -600.0 km per sec as the limit is to select stars with high radial velocities there by to select true hypervelocity stars (HVS) and not to confuse with high velocity stars and also it is based on the review paper of Brown (2015). Brown (2015) states that no one velocity threshold can define an unbound star, whether a star of a given speed is unbound or not depends on its location in the Milky Way. In this report I have not included HVS with postive radial velocities as it is a relatively lengthy paper it will not fit in this research note. They are reported in a separate paper (Parthasarathy 2023). The above mentioned search criteria has resulted eight HVS with accurate Gaia DR3 parallaxes. For these eight stars I derived their distances (D) from their parallaxes (Table 1). For the remainig seven HVS I used the distances given in Bailer-Jones et al. 2021) (see Table 1). I have derived for these 15 stars their U,V, W velocities and and tangential velocities (see Table 1). The U, V, W velocities also indicate they are hyper-velocity stars. The Teff, log g and [Fe/H] values of these stars (see also Gaia DR3 data of these stars from the SIMBAD database) indicate they are F-G stars. Gaia DR3 4094201527955913856 in this sample seems to be a very metal-poor HVS (Table 1). High resolution spectroscopy of these stars is needed to derive their atmospheric parameters, chemical composition and to understand their evolutionary status and origin.

Acknowledgements

This research has made use of the SIMBAD database, operated at CDS, Strasbourg, France

2000,A&AS,143,9, "The SIMBAD astronomical database", Wenger et al.

(see http://simbad.u-strasbg.fr/simbad/sim-basicIdent=m33&submit=SIMBAD+search)

TResearch breaching the Anthis paper is available for any on \$78274 esthightory. the American Astronomical Society There are no conflicts of interest. No funding from any source for this study.

Table 1. Fifteen hypervelocity stars detected from an analysis of Gaia DR3 data

Gaia DR3 No. I b parallax e mu-RA mu-DEC RUWE G D e

mas mas per year -- m pc 0 0 mas 4065480978657619968 7.3 -03.0 2.3202 0.0506 -6.172 -25.272 1.066 15.47 431 9 5305975869928712320 280 -03.3 0.3347 0.0212 -7.992 4.839 1.224 14.14 2987 189 5951114420631264640 342.3 -05.6 1.0612 0.0336 2.587 2.927 1.003 15.50 942 30 2198292118993038464 101.6 01.1 0.0459 0.0342 -2.027 -1.33 1.159 15.9 8560 4062883829092182144 2.8 -03.1 0.3817 0.1594 -5.240 -5.696 1.043 14.9 4101 5959019801816582272 347.9 -05.1 0.1639 0.0493 -2.74 -1.089 1.043 15.9 4844 4065791380145075072 6.0 -02.2 0.0526 0.0486 -2.967 -8.965 1.051 15.7 7671 4105689496051901440 21.4 -05.1 0.1032 0.0499 -1.227 -7.136 1.064 15.93 6187 5883971746674851840 325.1 -02.1 0.1372 0.0330 -6.282 -3.804 1.070 15.40 5677 1809832393157398016 57.4 -08.0 0.2211 0.0240 -2.759 -8.101 0.972 14.80 3898 5953456066818230528 345.5 -03.6 0.6990 0.0153 -5.720 -1.343 0.835 12.9 1431 31 4063258144073821184 3.96 -02.6 0.0927 0.0384 -2.618 -1.354 0.946 14.75 6826 1814359288672674560 64.5 - 13.5 0.1195 0.0115 - 0.840 - 8.031 0.914 12.8 8368 805 4100838558128545664 19.0 -07.7 0.1420 0.0269 -4.464 -3.465 1.043 15.1 5716 4094201527955913856 10.2 -0.86 0.2773 0.0275 -0.439 -1.561 1.126 14.59 3006 358

Radial Velocities, U, V, W and tangential velocities of HVS

Gaia DR3 No. RV e U V W Vtan

----- km per second km per second

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5305975869928712320 -	831 6	257 789 43 ´	132.3		
5951114420631264640 -	984 3	-605 772 86	17.5		
2198292118993038464 -	928 5	23 -926 -16	98.4		
4062883829092182144 -	909 1	-907 -40 45 1	50.5		
5959019801816582272 -	902 3	-615 656 73	67.7		
4065791380145075072 -	882 3	-878 -84 30 3	43.4		
4105689496051901440 -	737 10	-694 -242 59	212.3		
5883971746674851840 -	713 1	-274 658 23	197.6		
1809832393157398016 -	653 4	-402 -508 82	158.0		
5953456066818230528 -	634 5	-415 478 36	40.0		
4063258144073821184 -	631 1	-629 -39 26 9	95.4		
1814359288672674560 -	618 1	-320 -513 130	320.3		
4100838558128545664 -	601 6	-570 -176 72	153.0		
4094201527955913856 -	677 11	-668 9 91 2	23.1		
Teff, log g, [Fe/H] values of HVS					
Gaia DR3 No. Teff (K)	log g	[Fe/H] Com	ment		
4065480978657619968	4515 4.	574 +0.010			
5305975869928712320	5256 2.9	913 -0.790	Gaia Spectrum		
5951114420631264640	5356 4.	548 -0.328			
2198292118993038464	7123 2.	565 -1.079			
4062883829092182144		Gaia	light curves		
5959019801816582272	4842				

4065791380145075072 4	1933	2.658	0.034	
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4105689496051901440 4981 3.519 -0.056 ------

5883971746674851840 4858 2.304 -0.428 ------

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1809832393157398016	4498	2.154	0.019	Gaia Spectrum
5953456066818230528	4837	2.842	-0.053	Gaia Spectrum
4063258144073821184	4739	2.435	-0.191	
1814359288672674560			- Gaia	a Spectrum
4100838558128545664	5024	2.915	-0.100	
4094201527955913856	8400	2.854	-1.762	Gaia Spectrum, light curve

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