# Zodatianal Obsevatory. 

## BULLETIN No. CII.

## SUMMARY OF PROMINENCE OBSERVATIONS FOR THE SECOND HALF OF THE YEAR 1932

In pursuance of the programme of work adopted since lst January 1923 under the auspices of the International Astronomıcal Unıon, all observatorles takıng spectrohelograms of the sun have been asked to co-operate with the Kodarkanal Observatory by supplying copies of their photographs on those days when the Kodalkanal records are imperfect or wanting. In response to our requirements for the second half of the year 1932, the Mount Wilson Observatory supplied calcium ( $\mathrm{K}_{39}$ ) prominences plates for 55 days, H $a$ disc plates for 37 days and the Meuilon Obsorvatory supplied calcium ( $\mathrm{K}_{3}$ ) dise plates for 6 days and H $a$ dise plates for 36 days.

When only incomplete or imperfect photographs for any day are available from more than one observatory, the best photograph is chosen as representing the solar activity of that day, after weighting it according to its quality, and the remaining photographs are ignored.

Calcium Prominences at the Tımb.--The mean daily areas and numbers of prominences photographed during the half-year by means of the $\mathbb{K}$ line of calcium are given below. The means are corrected for meomplete or imperfect obscrvations, the total of 178 days for which plates were avalable being reduced to 165 effective days

Mean dally areas $\quad$| Mean danly |
| :---: |
| numbers. |

Compared with the previous half-year, areas and numbers show a decrease of 33 per cent and 22 per cent, respectively.

For comparison with bulletins issued prior to the co-operation of other observatories, the means based on Kodaikanal photographs alone are also given, 143 days of observation being counted as $122 \frac{1}{2}$ effective days.

|  |  |  |  | Mean dally areas (square mınutes) | Mean darly numbers |
| :---: | :---: | :---: | :---: | :---: | :---: |
| North (Kodaikanal photographs only) | . | .. | . | 0.98 | 4.50 |
| South ( do. ) | ... | - | ... | $0 \times 77$ | 424 |
|  |  | Total | ... | 175 | 8.74 |

The distribation of prominences in latitade is represented in the followng diagram, in which the full line gives the mean daily areas and the broken line the mean danly numbers for each zone of $5^{\circ}$ of latutude The ordnnates represent tenths of a square minute of are for the full line and numbers for the broken line Compared with the previous half year, the distribution of activity exhibits some well marked dufferences The peak near $30^{\circ} \mathrm{N}$ whech was seen in the first half of the year has now disappeared from the northern hemisphere and is evadenced in the southern hemisphere where the activity near $45^{\circ}$ has been much reduced


The monthly, quarterly and half-yearly areas and numbers and the mean height and mean extent of the prominences on photographs from all co operating observatories are given in Table I The unit of area is 1 square munute of arc The mean height is derived by adding together the greatest heights reached by individual prominences and dividing by the total number of prominences observed, the mean extent is derived by adding tegether the lengths of the base on the chromosphere of individual prominences and dividing by the total number of prominences -

Table I-Abstradt for the Shoond Half of 1932

| Months, | Number of days ( 0 eflectution). | Areas | Numbers | Daply means |  | Mean height | Mean extent |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | Areas | Numbers |  |  |
| 1982 |  |  |  |  |  | " | - |
| July | 29 | 533 | 262 | 18 | 90 | 304 | 359 |
| Angust | 28 | 452 | 213 | 16 | 76 | 351 | 437 |
| September - | 297 | 506 | 230 | 17 | 80 | 351 | 333 |
| October | 284 | 464 | 220 | 16 | 78 | 338 | 359 |
| November | 254 | 410 | 216 | 16 | 84 | 33 \% | 306 |
| December | 244 | 483 | 244 | 20 | 98 | 323 | 332 |
| Thurd quarter | 864 | 1491 | 705 | 17 | 82 | 334 | 374 |
| Fourth quarter | 789 | 1957 | 680 | 17 | 86 | 333 | 332 |
| Second half year | 165 | 2848 | 1,385 | 17 | 84 | 333 | 354 |

Distribution East and West of the Sun's Axus.-Compared with the previous half-year, areas showed a slight defect and numbers an excess at the east limb as will be seen from the following table -


Hydrogen Prominences at the Limb.-During the half-year, photographs of the prominences in hydrogen light were taken at this observatory on 101 days which were counted as 75 effective days The mean dauly areas of hydrogen prominences in square minutes of arc, are given below -

| Mean dally areas |  |  |  |  |  |  |  |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| North |  |  |  |  |  |  |  |  |
| (square minutes) |  |  |  |  |  |  |  |  |

Compared with the previous half-year, $\mathrm{H} a$ prominence areas show a decrease of 39 per cent. The ratio of $\mathrm{H} a$ areas to calcium areas is 35 per cent The curve of $\mathrm{H} a$ prommences is intermedrate between those of calcium prominences and $\mathrm{H} \imath \iota$ absorption markings.

Metalluc Prominences.-There were no metallic prominences observed during the half-year
Displacements of the Hydrogen Line.-Particulars of the displacements observed in the chromosphere and prominences are given in the following table.-
the
Table II.-Displacementsof/Hydrogen Line $/$.

| Date |  | $\begin{aligned} & \text { Time } \\ & \text { IS,T. } \end{aligned}$ |  | Latitude. |  | Displacement |  |  |  | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | North. | South | Limb | Red. | Violet | Both ways. |  |
| 1932 |  |  |  |  | M. | - | 。 |  | A | A | - A |  |
| July | 27 | 9 | 57 |  | 2 | $\underset{\text { E }}{ }$ |  | Slight |  | At top. |
|  | 28 |  | 53 |  | 25 | W |  |  |  |  |
| August | 13 |  | 40 |  | 475 | E |  | 05 |  | At top |
|  | 14 |  | 25 | 83 |  | E | $0 \cdot 5$ |  |  | In chromosphere |
|  | 15 |  | 13 |  | 26 | E |  | Slight |  | Do |
|  | 25 |  | 10 | 245 |  | E | 15 |  |  | At top |
|  |  |  | 04 |  | 24 | W |  | 05 |  | In chromosphere |
| September | 15 | 8 | 47 | 52.5 |  | W | 05 |  |  | Do |
|  | 21 | 8 | 35 | 80 |  | W | 05 |  |  | Att top. |
|  | 23 | 8 | 30 |  | 62 | W | 05 |  |  | Do |
|  |  | 8 | 20 |  | 2 | W |  | 1 |  | At base |
|  | 24 | 8 | 40 |  | 785 | W |  | 1 |  | At top |
| October | 20 | 9 | 45 | 2 |  | W | 15 |  |  | Do |
| November | 9 | 11 | 15 |  | ${ }_{80}{ }^{7}$ | $\underset{\text { E }}{\text { E }}$ | 1 |  |  | Do |
|  | 16 | 9 | 0 |  | 775 | $\underset{\text { E }}{\text { E }}$ |  | Slight |  | Do |
|  | 17 30 | 8 | 58 56 | 31 | 33 | $\stackrel{\text { E }}{\text { W }}$ | 05 | 1 |  | Do |
|  | 5 | 9 | 02 | 485 |  | E |  | 05 |  | Do |
| December | 5 | 9 | 07 | 215 |  | E | 1 |  |  | In chromosphere |
|  | 8 | 9 | 08 | 58 |  | E | 05 |  |  | At base |
|  | 18 | 8 | 55 | 145 |  | W | 2 |  |  | At top 15 m Displaced 25 A to Red at 9 h |
|  | 21 | 9 | $\bigcirc$ | 8 |  | W | Slight |  |  | At top |
|  | 27 | 8 | 39 |  | 28 | W | 05 |  |  | In chromosphere. <br> At top |
|  | 28 | 9 | 16 |  | 7 | $\underset{\mathbf{W}}{\mathbf{E}}$ |  | 05 |  | At top <br> At top. Extends over $2^{\circ}$ from $+14^{\circ}$ to |
|  |  | 9 | 27 | 15 |  | $\stackrel{W}{\mathbf{W}}$ | 1 |  |  | At top. Extends over $2^{\circ}$ from $+14^{\circ}$ to $+16^{\circ}$. |

The total number of displacements was 25 as against 80 in the previous half-year and their distribition was as follows -

|  |  | North | Sorth |
| :---: | :---: | :---: | :---: |
| $1^{\circ}-30^{\circ}$ |  | 6 | 7 |
| $31^{\circ}-60^{\circ}$ |  | 4 | 2 |
| $61^{\circ}-90^{\circ}$ |  | 2 | 4 |
|  |  | - | - |
|  | Total | 12 | 13 |
|  |  | - | - |
| East lumb |  |  |  |
| West limb |  |  |  |
|  | Total |  |  |

Of the displacements, 14 were towards the red and 11 towards the violet
Reversals and Dhsplacements on the Sun's Disc -Twenty-three bright reversals of the Halnne, 28 dark reversals of the $D_{s}$ line and 1 displacement of the $\mathrm{H} a$ lune were observed during the half-year Their disentm bution is given below -

|  | North | South | East | West |
| :--- | :---: | :---: | :---: | :---: |
| Bright reversals of $\mathrm{H} a$ | 16 | 7 | 11 | 12 |
| Dark raversals of $\mathrm{D}_{\mathbf{z}}$ | 15 | 8 | 12 | 11 |
| Displacements of $\mathrm{H} a$ | 1 |  | 1 |  |

The one displacement observed was towards the red
Prominences projected on the Dhsc as Absorption Markangs -Photographs of the sun's diso in Ha light were avalable from Kodarkanal and the co-operating observatories for a total of 176 days, which watik counted as 163 effective days The mean dally areas of $\mathrm{H} a$ absorption markings (corrected for foreshortening) in millionths of the sun's visible hemisphere and their mean daily nambers are given below -

|  |  | $\begin{aligned} & \text { Mean danly } \\ & \text { areas } \end{aligned}$ | Mean darly nambers |
| :---: | :---: | :---: | :---: |
| North |  | 553 | 388 |
| Sonth |  | 335 | 248 |
|  | Total | 888 | 636 |

The above show a decrease of 48 per cent in areas and 41 per cent in nambers, compared with the previous half-year
 on Kodaikanal photographs alone are also given, 120 days of observation being reckoned as 99 effective doditu

|  |  | Mean daly <br> areas | Mean dally <br> numbers |
| :---: | :---: | :---: | :---: |
| North (Kodaikanal photographs only) |  | 589 | 395 |
| Do | do | 292 | 217 |
|  | Total | -881 | $-\overline{612}$ |
|  |  | - | - |

The distribution of the mean dally areas in latitude is shown in the following diagram. The high latitude peak in the southern hemisphere is much reduced, while that in the belt $40^{\circ}-45^{\circ}$ in the northern hemisphere hass shifted $5^{\circ}$ towards the equator


Both areas and'numbers show an castern preponderance, the percentage east being 54 and 51 , lespectively, for areas and numbers. The arcas of $\mathrm{H} \alpha$ alsorption markings uncorrected for foreshortening are given below.-

|  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |  |  |  |  |
| Mean daily |  |  |  |  |  |  |  |  |  |  |
| areas |  |  |  |  |  |  |  |  |  |  |

The uncorrected areas amount to 50 per cent of the corrected ones as aganst 51 per cent for the previous hall-year.

The curve of distribution in latitude is sumilar to that for the corrected areas as usual.
Thanks are due to the co-operating observatories for the photographs supphed by them.

Kodatkanal, 13th September 1933.
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